

# Vacuum State and the Propagator of the Scalar Field in 3+1 Sph-Sym LQG

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# Outline

- 1 The 3+1 Spherically Symmetric Model
  - The Hamiltonian
  - Quantization and the Master Constraint
  - States
  - The Propagator

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# Canonical Transformation: Generic to Standard 3+1

- Standard 3+1 Hamiltonian (in Ashtekar variables)

$$\begin{aligned}
 H = N & \left( -\frac{E^\varphi}{2\sqrt{E^x}} - 2K_\varphi \sqrt{E^x} K_x - \frac{E^\varphi K_\varphi^2}{2\sqrt{E^x}} + \frac{((E^x)')^2}{8\sqrt{E^x} E^\varphi} - \frac{\sqrt{E^x} (E^x)' (E^\varphi)'}{2(E^\varphi)^2} \right. \\
 & \left. + \frac{\sqrt{E^x} (E^x)''}{2E^\varphi} + \frac{P_f^2}{2\sqrt{E^x} E^\varphi} + \frac{(E^x)^{3/2} (f')^2}{2E^\varphi} \right) \\
 & + N^1 \left( (E^x)' K_x - E^\varphi (K_\varphi)' - P_f f' \right)
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- $K_x$  and  $K_\varphi$ : related to radial and angular components of the extrinsic curvature of the spatial hypersurfaces.

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# Strategy to Attack the Problem of Quantization

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Algebra of constraints in general relativity is a **non-Lie algebra**:

The Poisson bracket of Hamiltonian constraint with itself contains **structure functions** rather than constants.

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## The problem of dynamics

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- The algebra of (smeared) Hamiltonian constraints does not close, it is proportional to a spatial diffeomorphism constraint.
- The coefficient of proportionality is not a constant, it is a non-trivial function on the phase space.

$$\{\vec{C}(\vec{N}), \vec{C}(\vec{N}')\} = \kappa \vec{C}(\mathcal{L}_N \vec{N}')$$

$$\{\vec{C}(\vec{N}), C(N')\} = \kappa C(\mathcal{L}_N \vec{N}')$$

$$\{C(N), C(M)\} = \int d^3x (N \partial_a M - M \partial_a N) g^{ab} C_b$$

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or its discrete version

$$\mathbb{H}^\epsilon = \sum_i \mathbb{H}(i) = \sum_i \frac{1}{2} \frac{\mathcal{H}(i)^2 \ell_P}{\sqrt{g(i)}} = \sum_i \frac{1}{2} \frac{\mathcal{H}(i)^2 \ell_P}{\sqrt{E^x(i)} E^\varphi(i)}$$

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- Claims that by using this:
  - Problems with the commuter algebra disappear.
  - Could have control of the solution space.
  - Could have control of the (quantum) Dirac observables of LQG.
  - Even a decision on whether the theory has the correct classical limit.
  - The connection with the path (or spin foam) formulation could be within reach.

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- Configuration variable(s) are written as

$$K_\varphi \rightarrow \frac{\sin(\rho K_\varphi)}{\rho}$$

and one assumes that the radial direction is bounded with a spatial extent  $L$  and consists of discrete points  $x_i$  separated by a coordinate distance  $\epsilon$ , thus e.g.

$$\int dx \rightarrow \epsilon \sum_x,$$

$$\delta(x-y) \rightarrow \frac{\delta_{x,y}}{\epsilon},$$

$$\frac{\delta}{\delta f(x)} \rightarrow \frac{1}{\epsilon} \frac{\partial}{\partial f},$$

$$f(x)' \rightarrow \frac{f(x_{i+1}) - f(x_i)}{\epsilon},$$



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- **Physical states**: cylindrical functions in the **kernel of  $\mathbb{H}$** , i.e.  $\mathbb{H}|\Psi_{\text{cyl}}\rangle = 0$ .
- Find this kernel by variational technique: **minimize the expectation value of master constraint**  $\langle \Psi_{\text{cyl}} | \mathbb{H} | \Psi_{\text{cyl}} \rangle$ .

# Quantized Master Constraint

- **Discretize** the Hamiltonian constraint  $H = H_{\text{vac}} + GH_{\text{matt}}$ , **holonomize the gravitational** degrees of freedom, **quantize the matter** degrees of freedom **using Fock** quantization (to be able to compare with QFT).

# Quantized Master Constraint

- Partially gauge fixed ( $E^x = x^2$ ), quantized Hamiltonian constraint reads

$$H(i) = - (1 - 2\Lambda)\epsilon - x(i+1) \frac{\sin^2(\rho K_\varphi(i+1))}{\rho^2} + x(i) \frac{\sin^2(\rho K_\varphi(i))}{\rho^2} + \frac{x(i+1)^3 \epsilon^2}{(E^\varphi(i+1))^2} - \frac{x(i)^3 \epsilon^2}{(E^\varphi(i))^2} + \ell_p^2 \left( \frac{H_{\text{matt}}^{(1)}(i)}{(E^\varphi)^2(i)} + \frac{H_{\text{matt}}^{(2)}(i) \sin(\rho K_\varphi(i))}{\rho E^\varphi(i)} - H_{\text{matt}}^{(3)}(i) \right)$$

with  $\Lambda = \frac{G}{2} \rho_{\text{vac}}$  and the vacuum energy density  $\rho_{\text{vac}}$  and

$$H_{\text{matt}}^{(1)}(i) = \left( \epsilon (P_f(i))^2 + \epsilon x(i)^4 (f(i+1) - f(i))^2 \right) \ell_p^2,$$

$$H_{\text{matt}}^{(2)}(i) = (-2x(i) (f(i+1) - f(i)) P_f(i)) \ell_p^2,$$

$$H_{\text{matt}}^{(3)}(i) = 2\rho_{\text{vac}} \epsilon \ell_p^2.$$

This way the expectation value of  $H_{\text{matt}}$  will be zero in the vacuum.

# Quantized Master Constraint

- The discrete master constraint *operator* reads

$$\begin{aligned} \hat{\mathbb{H}}(i) = \ell_P \left[ \hat{c}_{11}(i) \left( \hat{H}_{\text{matt}}^{(1)}(i) \right)^2 + \hat{c}_{22}(i) \left( \hat{H}_{\text{matt}}^{(2)}(i) \right)^2 + \hat{c}_{33}(i) \left( \hat{H}_{\text{matt}}^{(3)}(i) \right)^2 \right. \\ \left. + \hat{c}_1(i) \hat{H}_{\text{matt}}^{(1)}(i) + \hat{c}_2(i) \hat{H}_{\text{matt}}^{(2)}(i) + \hat{c}_3(i) \hat{H}_{\text{matt}}^{(1)}(i) + \hat{c}_{00}(i) \right. \\ \left. + \hat{c}_{12}(i) \hat{H}_{\text{matt}}^{(1)}(i) \hat{H}_{\text{matt}}^{(2)}(i) + \hat{c}_{13}(i) \hat{H}_{\text{matt}}^{(1)}(i) \hat{H}_{\text{matt}}^{(3)}(i) + \hat{c}_{23}(i) \hat{H}_{\text{matt}}^{(2)}(i) \hat{H}_{\text{matt}}^{(3)}(i) \right] \end{aligned}$$

the  $\hat{c}_i$  and  $\hat{c}_{ij}$  coefficients contain only gravitational degrees of freedom,  $K_\varphi$  and  $E^\varphi$ .

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# Construction of the States: Gravitational Part

- We are interested in the vacuum solutions: classically correspond to  $f = P_f = 0$ . Thus for now we ignore  $H_{\text{matt}}$  and only consider  $H_{\text{vac}}$ :

$$H_{\text{vac}} = \left( -x(1 - 2\Lambda) - xK_{\varphi}^2 + \frac{x^3}{(E_{\varphi})^2} \right)'$$

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- Make another gauge fixing  $K_\varphi = 0$ . Then weakly vanishing of  $H_{\text{vac}}$  means

$$E^\varphi = \frac{x}{\sqrt{1 - 2\Lambda}}$$

which is the classical solution.

# Construction of the States: Gravitational Part

- Construct a polymer representation: set up a lattice of points  $j = 0 \dots N$  in the radial direction and write a “point holonomy” for the  $K_\varphi$  variable at each lattice site,

$$T_{\vec{\mu}} = \exp \left( i \sum_j \mu_j K_\varphi(j) \right) = \langle K_\varphi | \vec{\mu} \rangle$$

The quantities  $\mu_i$  play the role of the “loop” in this one dimensional context.

# Construction of the States: Gravitational Part

- The trial quantum states of the gravitational part  $\langle \vec{\mu} | \psi_{\vec{\sigma}} \rangle$  are chosen to be centered around the classical solution

$$\langle \vec{\mu} | \psi_{\vec{\sigma}} \rangle = \prod_i \sqrt[4]{\frac{2}{\pi \sigma(i)}} \exp \left( -\frac{1}{\sigma(i)} \left( \mu_i - \frac{1}{\ell_p^2} \left( \frac{\epsilon x(i)}{\sqrt{1-2\Lambda}} \right) \right)^2 \right)$$

with the variable  $\mu_i$  to be centered around the classical value of

$$E^\varphi(i) = \frac{\epsilon x(i)}{\sqrt{1-2\Lambda}}.$$

# Construction of the States: Matter Part

- Find the expectation value of the matter Hamiltonian on the state  $\langle \vec{\mu} | \psi_{\vec{\sigma}} \rangle$ . This will be an operator  $\hat{H}_{\text{matt}}^{\text{eff}}$  acting on the matter fields.

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- Write  $\hat{H}_{\text{matt}}^{\text{eff}}$  in terms of creation-annihilation  $\hat{C}$  and  $\hat{\bar{C}}$  operators by using Fourier analysis. It is

$$\hat{H}_{\text{matt}}^{\text{eff}} = \langle \psi_{\vec{\sigma}} | \hat{H}_{\text{matt}} | \psi_{\vec{\sigma}} \rangle = (1 - 2\Lambda) \int_0^{2\pi/\epsilon} d\omega \omega \hat{C}(\omega) \hat{C}(\omega).$$

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- The vacuum states  $|0\rangle$  are the states of operator  $\hat{H}_{\text{matt}}^{\text{eff}}$ , annihilated by  $\hat{C}$ .

# The Full Trial State

- The full trial states: direct product of the vacuum of the matter part of the Hamiltonian and the Gaussian on the gravitational variables:

$$|\psi_{\vec{\sigma}}^{\text{trial}}\rangle = |\psi_{\vec{\sigma}}\rangle \otimes |0\rangle$$



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$$|\psi_{\vec{\sigma}}^{\text{trial}}\rangle = |\psi_{\vec{\sigma}}\rangle \otimes |0\rangle$$

- To find the physical states,  $\hat{\mathbb{H}}|\psi_{\vec{\sigma}}^{\text{trial}}\rangle_{\text{phys}} = 0$ , vary  $\vec{\sigma}$  to find the minimum of the expectation value of the master constraint on the trial states  $\langle\psi_{\vec{\sigma}}^{\text{trial}}|\hat{\mathbb{H}}|\psi_{\vec{\sigma}}^{\text{trial}}\rangle$ .

# Minimizing the Master Constraint

- Minimizing  $\langle \psi_{\vec{\sigma}}^{\text{trial}} | \hat{\mathbb{H}} | \psi_{\vec{\sigma}}^{\text{trial}} \rangle$  results in

$$\langle \psi_{\vec{\sigma}}^{\text{trial}} | \hat{\mathbb{H}} | \psi_{\vec{\sigma}}^{\text{trial}} \rangle = \frac{\sigma_0 \ell_{\text{P}}^3}{\epsilon x^2} + \mathcal{O}(\ell_{\text{P}}^5)$$

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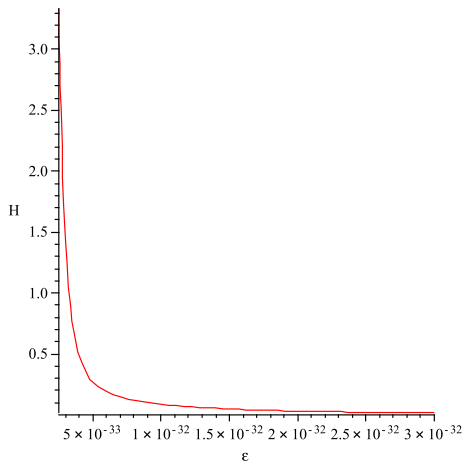
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with  $\sigma_0$  of order unity.

- Also we have assumed  $\sigma$  independent of  $x$ . Our experiments suggest that allowing variations in  $x$  leads to the same minimum value of  $\sigma$  approximately independent of  $x$ .

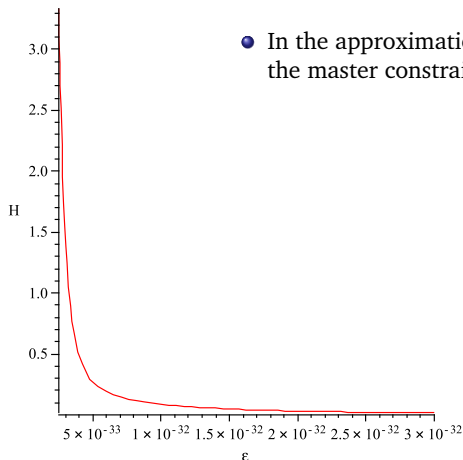
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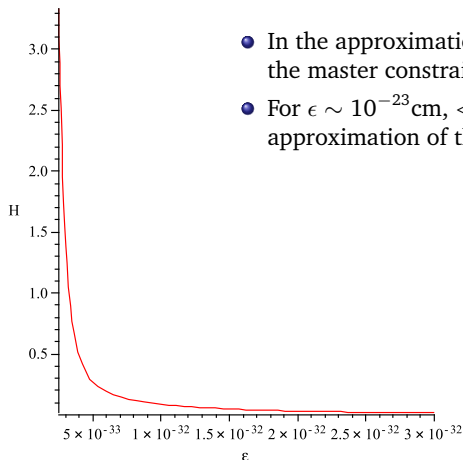
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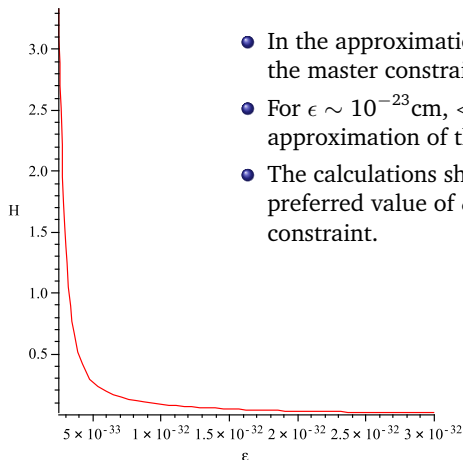
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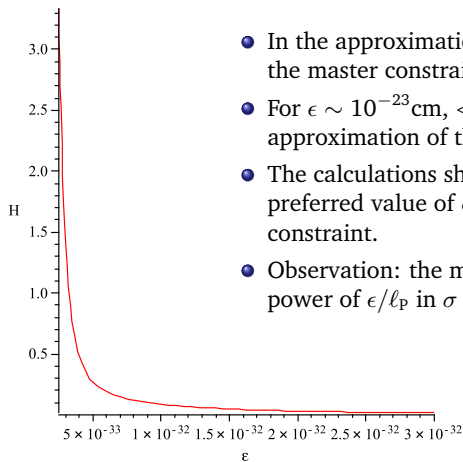


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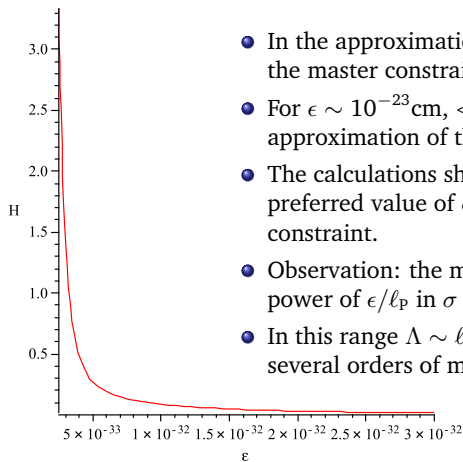
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- In this range  $\Lambda \sim \ell_p^2/\epsilon^2$  would not be of Planck scale but several orders of magnitude smaller.

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- Is it appropriate to use this vacuum in the state  $|\psi_{\vec{\sigma}}^{\text{trial}}\rangle$  to compute the propagator?
- To answer: show that the corrections to the Fock representation due to holonomization are small enough for our purpose.

# Is Fock Vacuum Appropriate?

- Fully holonomized Hamiltonian

$$H_{\text{matt}}(i) = \frac{H^{(1)}(i)}{(E^\varphi(i))^2} + \frac{H^{(2)}(i) \sin(\rho K_\varphi(i))}{\rho E^\varphi(i)},$$

where,

$$H^{(1)}(i) = \frac{\epsilon}{2} P_f^2(i) x(i)^2 + \frac{\epsilon^3 \sin^2(\beta f(i))}{2\beta^2} - \frac{\epsilon^2 x(i)}{\beta^2} \sin[\beta f(i) \sin(\beta(f(i+1) - f(i)))] \\ + \frac{\epsilon x(i)^2}{2\beta^2} \sin^2(\beta(f(i+1) - f(i))),$$

$$H^{(2)}(i) = \frac{P_f(i)}{\beta} (\epsilon \sin(\beta f(i)) - x(i) \sin(\beta(f(i+1) - f(i))))$$

with  $\beta$  being the holonomization parameter.

# Is Fock Vacuum Appropriate?

- Expanding in  $\beta$  and keeping the two lowest orders:

$$H^{(1)}(i) = H_{\text{lead}}^{(1)}(i) + H_{\text{corr}}^{(1)}(i),$$

$$H^{(2)}(i) = H_{\text{lead}}^{(2)}(i) + H_{\text{corr}}^{(2)}(i)$$

leading orders correspond to the Fock terms; correction terms of order  $\beta^2$  are holonomization corrections.

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- The correction part of the master constraint:

$$\begin{aligned} \mathbb{H}_{\text{corr}}(i) = & c_1(i)H_{\text{corr}}^{(1)}(i) + c_2(i)H_{\text{corr}}^{(2)}(i) + c_{00}(i) + c_{11}(i) \left( H_{\text{lead}}^{(1)}(i)H_{\text{corr}}^{(1)}(i) \right) \\ & + c_{12}(i) \left( H_{\text{lead}}^{(1)}(i)H_{\text{corr}}^{(2)}(i) + H_{\text{lead}}^{(2)}(i)H_{\text{corr}}^{(1)}(i) \right) + c_{22}(i) \left( H_{\text{lead}}^{(2)}(i)H_{\text{corr}}^{(2)}(i) \right) \end{aligned}$$



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- Show that  $\int dx \langle \psi_{\vec{\sigma}}^{\text{trial}} | \mathbb{H}_{\text{corr}} | \psi_{\vec{\sigma}}^{\text{trial}} \rangle$  is very small.

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- Thus for  $\epsilon \gg \ell_p \Rightarrow \mathbf{H}_{\text{corr}} \ll \mathbf{H}_{\text{lead}}$ : the vacuum of the theory is well approximated by the tensor product  $|\psi_{\vec{\sigma}}^{\text{trial}}\rangle = |\psi_{\vec{\sigma}}\rangle \otimes |0\rangle$ : we can continue to use it to compute the propagator.

# The Choices of Polymerization

- The discretized Hamiltonian written in a more symmetric way

$$H = \sum_i \frac{P_f(i)^2}{2\epsilon} - \frac{(f(i+1) + f(i-1) - 2f(i))f(i)}{2\epsilon}$$

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- Several choices for polymerization.
- We go for two choices: the field  $f$  itself; the momentum  $P_f$ . They do not lead to equivalent theories. Also polymerizing the momentum yields a theory that in the continuum limit is not polymeric.

# Polymerizing the Scalar Field

- The polymerized Hamiltonian in this case

$$H = \sum_i \left( \frac{P_f(i)^2}{2\epsilon} - \frac{\sin(\beta(f(i+1) + f(i-1) - 2f(i))) \sin(\beta f(i))}{2\epsilon\beta^2} \right)$$



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- Perturbative expansion in  $\beta$  yields  $H = H_0 + H_{\text{int}}$  with

$$H_0 = \sum_i \left( \frac{P_f(i)^2}{2\epsilon} - \frac{f(i)(f(i+1) + f(i-1) - 2f(i))}{2\epsilon} \right)$$

$$H_{\text{int}} = \sum_i \frac{1}{2\epsilon} \left[ \frac{1}{6} f(i)(f(i+1) + f(i-1) - 2f(i))^3 \right. \\ \left. + \frac{1}{6} f(i)^3 (f(i+1) + f(i-1) - 2f(i)) \right] \beta^2$$

# Polymerizing the Scalar Field

- The propagator to leading order is

$$G^{(2)}(j, t, k, t') = G^{(0)}(j, t, k, t') + \frac{i^2}{2!} \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{\infty} dt_2 \\ \times \sum_{j'=-N}^N \sum_{k'=-N}^N \langle 0 | T (f(j, t) f(k, t') H_{\text{int}}(j', t_1) H_{\text{int}}(k', t_2)) | 0 \rangle$$

where we use  $a, \dots, j, k$  for direct space and  $m, n, \dots, z$  for momentum space.

# Polymerizing the Scalar Field

- Going from direct to momentum representation ( $\dots, j, k \rightarrow m, n, \dots$ ) and from time  $t$  to  $\omega$  space, with  $p(n) \equiv \pi n/L$  and  $L = N\epsilon$  yields

$$\begin{aligned}
 G^{(2)}(n_1, \omega_1, n_2, \omega_2) &= G^{(0)}(n_1, \omega_1, n_2, \omega_2) \\
 &+ \left[ \frac{\alpha_1 \beta^4}{\epsilon^2} + \beta^4 \alpha_2 p(n_1)^2 \right] \frac{4\pi i}{\epsilon} \frac{\delta(\omega_1 - \omega_2) (\delta_{n_1, n_2} - \delta_{n_1, -n_2})}{(\omega_1^2 - p(n_1)^2 + i\sigma)^2} \\
 &\approx \frac{4\pi i}{\epsilon} \frac{1}{\omega_1^2 - p(n_1)^2 (1 + \alpha_2 \beta^4) - \frac{\alpha_1 \beta^4}{\epsilon^2} + i\sigma} \\
 &\times (\delta_{n_1, n_2} - \delta_{n_1, -n_2}) \delta(\omega_1 - \omega_2)
 \end{aligned}$$

to lowest order in  $p(n)$ . With  $\alpha_1$  and  $\alpha_2$  constants of order one.

# Polymerizing the Momentum Field

- The same method of calculations for the polymerizing the momentum in the Hamiltonian

$$H = \sum_i \frac{\sin^2(\beta P_f(i))}{2\beta^2 \epsilon} - \frac{(f(i+1) + f(i-1) - 2f(i))f(i)}{2\epsilon}$$

yields

$$\begin{aligned} G^{(2)}(n_1, \omega_1, n_2, \omega_2) &= G^{(0)}(n_1, \omega_1, n_2, \omega_2) \\ &\quad + \beta^4 \alpha_2 p(n_1)^2 \frac{4\pi i}{\epsilon} \frac{\delta(\omega_1 - \omega_2) (\delta_{n_1, n_2} - \delta_{n_1, -n_2})}{(\omega_1^2 - p(n_1)^2 + i\sigma)^2} \\ &\approx \frac{4\pi i}{\epsilon} \frac{1}{\omega_1^2 - p(n_1)^2 (1 + \alpha_2 \beta^4) + i\sigma} \\ &\quad \times (\delta_{n_1, n_2} - \delta_{n_1, -n_2}) \delta(\omega_1 - \omega_2) \end{aligned}$$

to lowest order in  $p(n)$ . With  $\alpha_2$  constant of order one.

# Lorentz Invariance Violation

- In both of the propagators we derived, generically there will be higher powers of  $p$  in the denominator (we only kept the lowest). They are in the class of Hořava's "Gravity at the Lifshitz point":

$$\frac{1}{\omega^2 - c^2 \mathbf{k}^2 - G(\mathbf{k}^2)^z}$$

which signal Lorentz invariance violation.

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## Corrections due to Polymerization

The order in  $\beta$  at which corrections appear can be shifted (and made small) arbitrarily by choosing suitable polymerizations of the theory: e.g. the corrections to the dispersion relation can be made to be of order  $\beta^8$  instead of  $\beta^4$ , etc.

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- The form of the Lorentz violation depends on how one polymerizes.

# References

This work is done under supervision of Rodolfo Gambini and in collaboration with Jorge Pullin (LSU). For more details see:

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