

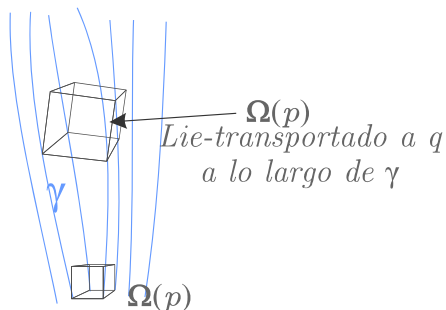
# The Raychaudhuri Equation for the Interior of Effective Loop Quantum Black Holes

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## Evolution of spatial volume



Raychaudhuri identity:

$$\frac{d\theta}{d\tau} = -\frac{1}{3} \underbrace{\theta^2}_{\text{expansion}} - \underbrace{\sigma^{ab}\sigma_{ab}}_{\text{shear}} + \underbrace{\omega^{ab}\omega_{ab}}_{\text{twist}} - \underbrace{R_{ab}\gamma^a}_{\text{Ricci}} \underbrace{\gamma^b}_{\text{tangent to } \gamma}$$

Evolution of  $\Omega$  :

$$\mathcal{L}_\gamma \Omega = \theta \Omega \iff \{H, \Omega\} = \theta \Omega$$

## Expansion factor and singularities.

- Attractive or Repulsive interaction?

$$\theta > 0 \quad \text{or} \quad \theta < 0$$

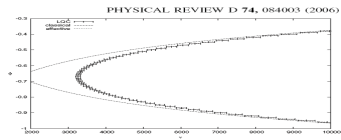
- Classical GR + the strong energy condition ( $R_{ab}\gamma^a\gamma^b \geq 0$ ) implies:

$$\theta \rightarrow -\infty$$

- Divergence of  $\theta$  is a necessary condition in singularity theorems

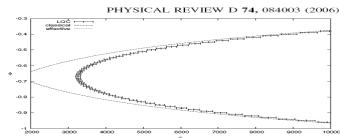
# Friedmann-Robertson-Walker singularity resolution

- Ashtekar, Pawłowski and Singh:



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- $\theta$  is bounded:

$$H_{\text{eff}}^2 = \frac{\theta_{\text{eff}}^2}{9} = \frac{1}{4} \frac{1}{\lambda^2 \gamma} \sin^2(2\lambda\beta) \approx \frac{8\pi G}{3} \rho \left(1 - \frac{\rho}{\rho_{\text{crit}}}\right) \xrightarrow{\rho \ll \rho_{\text{crit}}} \frac{8\pi G}{3} \rho$$

$\downarrow$   $4\sqrt{3}\pi\gamma l_{\text{Pl}}^2$  Immirzi       $\downarrow$  canonical coordinate       $\downarrow$   $0.41\rho_{\text{Planck}}$

$$\rho \approx \frac{3}{8\pi G\lambda^2} \frac{\sin^2(\lambda\beta)}{\lambda^2}$$

# Black Hole Interior (Kantowski-Sachs)

- Symmetry group:  $\mathbb{R} \times SU(2)$

$$[\tilde{\zeta}_i, \tilde{\zeta}_j] = \epsilon_{ij}^k \underset{\in \mathfrak{su}(2)}{\tilde{\zeta}_k} \quad [\tilde{\zeta}_i, \zeta] = 0$$

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- + No matter
- + Einstein Field Equations

$\implies$  Schwarzschild Interior:

$$ds^2 = - \left( 1 - \frac{2m}{T} \right) dR^2 + \frac{1}{\left( 1 - \frac{2m}{T} \right)} dT^2 + T^2 d\Omega^2$$

# Hamiltonian formulation in Ashtekar variables

- Kantowski-Sachs ( $\mathbb{R} \times SO(3)$ ) symmetric connection and triads:

$$A = A_a^i \tau_i dx^a = L^{-1} c \tau_3 dx + b \tau_2 d\theta - b \tau_1 \sin \theta d\phi + \tau_3 \cos \theta d\phi,$$

$\downarrow$   
 $-i\sigma_i/2$

$$\tilde{E} = p_c \tau_3 \sin \theta \partial_x + L^{-1} p_b \tau_2 \sin \theta \partial_\theta - L^{-1} p_b \tau_1 \partial_\phi, \quad (1)$$

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$$qq^{ab} = \delta^{ij} \tilde{E}^a_i \tilde{E}^b_j, \quad (3)$$

$$ds^2 = -N(t)^2 dt^2 + g_{xx}(t) dx^2 + g_{\theta\theta}(t) d\theta^2 + g_{\phi\phi}(t) d\phi^2, \quad (4)$$

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- Geometrical quantities

$$S^2 - \text{area: } A = 4\pi p_c; \quad (I \times S^2) - \text{volume: } \Omega = 4\pi p_b \sqrt{p_c}. \quad (6)$$

$$\text{Expansion Factor: } \theta = \frac{\dot{\Omega}}{\Omega} \quad (7)$$

$$\text{Shear: } \sigma = \frac{1}{\sqrt{3}} \left( \frac{1}{p_b} \dot{p}_b - \frac{1}{p_c} \dot{p}_c \right) \quad (8)$$

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- Hamiltonian "al gusto"



$$H_{cl} = -\frac{N}{2G\gamma^2} \left[ 2bc\sqrt{p_c} + (b^2 + \gamma^2) \frac{p_b}{\sqrt{p_c}} \right].$$

# Classical dynamics



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$$\theta = \frac{1}{\Omega} \{ \Omega, H[N=1] \} = \frac{2}{\gamma} \left( 2 \frac{b}{\sqrt{p_c}} + \frac{c\sqrt{p_c}}{p_b} \right)$$

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- On the constraint surface:

$$\theta \approx \frac{1}{\gamma} \frac{1}{b \sqrt{p_c}} (3b^2 - \gamma^2)$$

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$$\begin{aligned} R_{00} &= -\frac{1}{3}\theta^2 - \sigma^{ab}\sigma_{ab} + \omega^{ab}\omega_{ab} - \frac{d\theta}{d\tau} \\ &= \frac{1}{2\gamma^2 p_b p_c} (b^2 p_b + 2bc p_c + p_b^2) \\ &\approx 0 \rightarrow \text{Einstein Field Eq. in vac} \end{aligned}$$

# Effective Semiclassical Black Hole InteriorS

- "Holonimize" classical connection::

$$b \longrightarrow \frac{\sin(\mu_b b)}{\mu_b}, \quad c \longrightarrow \frac{\sin(\mu_c c)}{\mu_c}$$

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$\mu_0$ -scheme:

$$\mu_b = cte, \quad \mu_c = cte$$

$\bar{\mu}$ -scheme:

$$\bar{\mu}_b = \sqrt{\frac{\Delta}{p_b}}, \quad \bar{\mu}_c = \sqrt{\frac{\Delta}{p_c}},$$

Very succesfull  $\bar{\mu}'$ -scheme:

$$\bar{\mu}'_b = \sqrt{\frac{\Delta}{p_c}}, \quad \bar{\mu}'_c = \frac{\sqrt{p_c \Delta}}{p_b}.$$

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- Hamiltonian

$$H_{eff} = -\frac{N}{2G\gamma^2} \left[ 2 \frac{\sin(\mu_b b)}{\mu_b} \frac{\sin(\mu_c c)}{\mu_c} \sqrt{p_c} + \left( \frac{\sin(\mu_b b)}{\mu_b} \right)^2 + \gamma^2 \right] \frac{p_b}{\sqrt{p_c}}$$

# Effective expansion factors

- $\mu_0$  :

$$\begin{aligned}\theta_0 = & + \frac{1}{\rho_b} \frac{1}{\gamma} \sqrt{\rho_c} \frac{\sin(\mu_c c)}{\mu_c} \cos(\mu_b b) + \frac{1}{\sqrt{\rho_c}} \frac{1}{\gamma} \frac{\sin(\mu_b b)}{\mu_b} \cos(\mu_c c) \\ & + \frac{1}{\sqrt{\rho_c}} \frac{1}{\gamma} \frac{\sin(\mu_b b)}{\mu_b} \cos(\mu_c c)\end{aligned}$$

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- $\bar{\mu}$  :

$$\begin{aligned}\bar{\theta} = & + \frac{1}{\rho_b} \frac{1}{\gamma} \frac{1}{\sqrt{\Delta}} \rho_c \sin\left(\sqrt{\frac{\Delta}{\rho_c}} c\right) \cos\left(\sqrt{\frac{\Delta}{\rho_b}} b\right) \\ & + \frac{1}{\sqrt{\rho_c}} \frac{1}{\gamma} \frac{1}{\sqrt{\Delta}} \rho_b^{\frac{1}{2}} \sin\left(\sqrt{\frac{\Delta}{\rho_b}} b\right) \cos\left(\sqrt{\frac{\Delta}{\rho_b}} b\right) \\ & + \frac{1}{\sqrt{\rho_c}} \frac{1}{\gamma} \frac{1}{\sqrt{\Delta}} \rho_b^{\frac{1}{2}} \sin\left(\sqrt{\frac{\Delta}{\rho_b}} b\right) \cos\left(\sqrt{\frac{\Delta}{\rho_c}} c\right)\end{aligned}$$

# Effective gravitational repulsion

- $\bar{\mu}'$

$$\begin{aligned}\bar{\theta}' &= \frac{1}{\Omega} [\Omega, H_{eff}] \\ &= \frac{1}{\gamma} \frac{1}{\sqrt{\Delta}} \left[ \sin \left( \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right) \cos \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) + \right. \\ &\quad \left. + \sin \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) \cos \left( \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right) + \sin \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) \cos \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) \right]\end{aligned}$$

$$\theta_{fis} = \frac{\rho_c \sin(b\mu_b) \left( \sqrt{-\frac{\gamma^4 \Delta^2 \csc^2(b\mu_b)}{\rho_c^2} - \sin^2(b\mu_b) - \frac{2\gamma^2 \Delta}{\rho_c} + 4 + \cos(b\mu_b)} \right) - \gamma^2 \Delta \cot(b\mu_b)}{2\gamma\sqrt{\Delta}\rho_c}$$

*BOUNDED ON ALL PHASE SPACE*

*Gravitational repulsion*

*without energy conditions violations*



## Shear is bounded:

$$\begin{aligned}\bar{\sigma}' &= \frac{1}{\sqrt{3}} \left( \frac{1}{\rho_b} \dot{\rho}_b - \frac{1}{\rho_c} \dot{\rho}_c \right) \\ &= \frac{1}{\sqrt{3}} \frac{1}{\gamma} \frac{1}{\sqrt{\Delta}} \left( \sin \left( \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right) \cos \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) \right. \\ &\quad \left. + \sin \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) - 2 \sin \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) \cos \left( \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right) \right)\end{aligned}$$

# Expansion factor evolution

•  
 $\dot{\theta} =$  (bounded terms on all phase space)

+ (bounded terms on the constraint surface)

$$+ \frac{1}{\gamma^2} \frac{1}{\Delta} \left[ \frac{\sqrt{\rho_c \Delta}}{\rho_b} c - \sqrt{\frac{\Delta}{\rho_c}} b \right] \cos \left( \sqrt{\frac{\Delta}{\rho_c}} b - \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right) \sin \left( \sqrt{\frac{\Delta}{\rho_c}} b - \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right)$$

$$+ \frac{1}{\gamma^2} \frac{1}{\Delta} \left( \frac{\sqrt{\rho_c \Delta}}{\rho_b} c - \sqrt{\frac{\Delta}{\rho_c}} b \right) \cos \left( 2 \sqrt{\frac{\Delta}{\rho_c}} b \right) \sin \left( \sqrt{\frac{\Delta}{\rho_c}} b \right) \cos \left( \frac{\sqrt{\rho_c \Delta}}{\rho_b} c \right)$$

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• Need to know something about  $\left( \sqrt{\frac{\Delta}{p_c}} b - \frac{\sqrt{p_c \Delta}}{p_b} c \right)$

• Qualitative analysis+theory of differential equations+numerical investigations

## Qualitative analysis of the constraint

$$H = 2 \sin \left( \sqrt{\frac{\Delta}{p_c}} b \right) \sin \left( \frac{\sqrt{p_c \Delta}}{p_b} c \right) + \sin^2 \left( \sqrt{\frac{\Delta}{p_c}} b \right) + \frac{\Delta \gamma^2}{p_c} \approx 0$$

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- For finite  $p_c$ ,  $p_b$  is bounded above onshell.

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- $\rho_c$  is bounded below on the constraint surface:

$$\Delta \gamma^2 \lesssim |\rho_c|.$$

- $b = 0$  is not on the constraint surface
- $c = 0$  is not on the constraint surface, unless  $\rho_c \rightarrow \infty$
- For finite  $\rho_c$ ,  $\rho_b$  is bounded above onshell.
- As  $\rho_c$  is bounded below and  $\rho_b \neq 0$ ,  $\mathbf{V} = 4\pi\rho_b\sqrt{\rho_c} \neq \mathbf{0}$ . (Classically  $\mathbf{V} \rightarrow \mathbf{0}$  on the physical singularity and on the horizon).

Numerical analysis is on work...

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*THANKS!*